



*Beyond Einstein: From the Big Bang to Black Holes*

## *Frequency-Tunable Pre-stabilized lasers for LISA via Sideband Locking*

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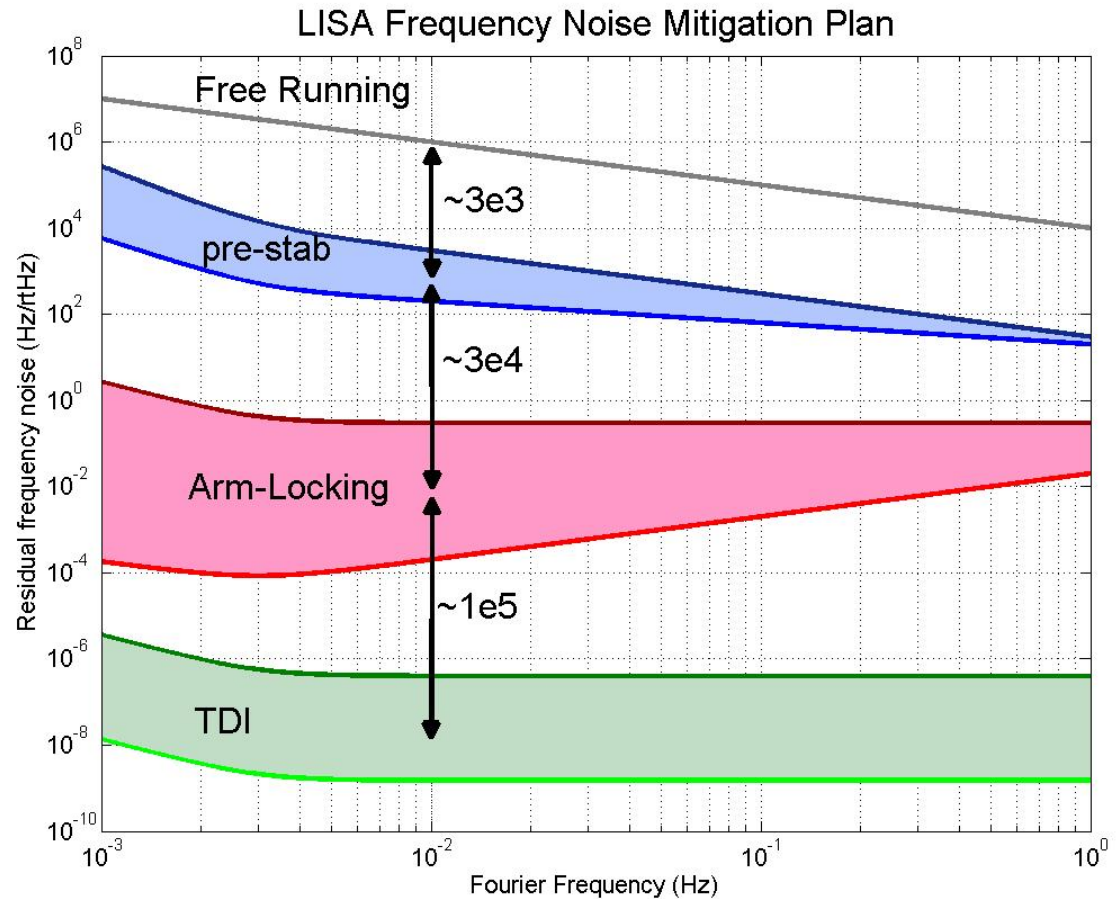
7<sup>th</sup> International LISA Symposium

June 18<sup>th</sup>, 2008

Barcelona, Spain

## Laser frequency noise is a major potential noise source for LISA

- Three-stage system (two active one passive) to achieve overall suppression of  $\sim 10^{13}$
- Running pre-stabilization and arm-locking in series reduces gain (bandwidth) requirements on arm-locking.
- Serial arrangement *requires frequency-tunable pre-stabilization* because cavity lock point drifts with respect to arm-locking lock point.



## Tunable Cavity

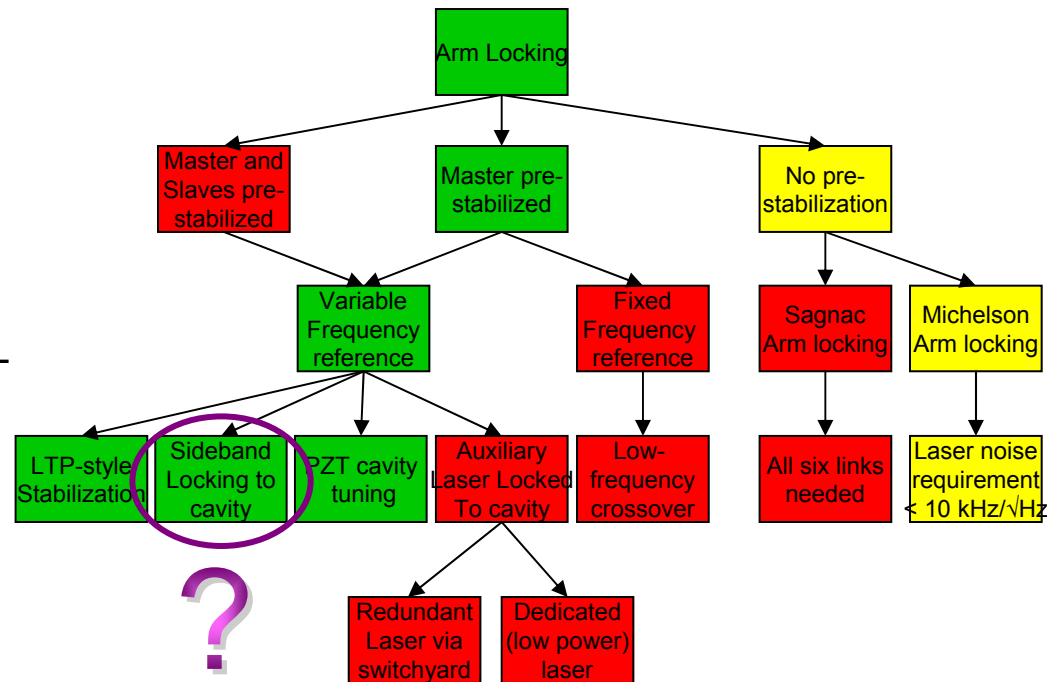
- Adjust length of cavity spacer
- “Standard” method
- Likely reduces cavity stability

## Offset Phase-Lock Loop

- Lock master laser to cavity, PLL slave with adjustable offset
- No modification to cavity
- Requires additional laser

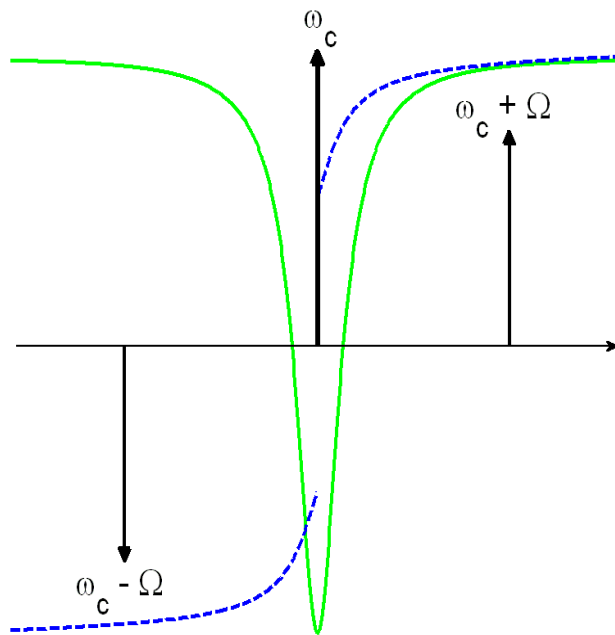
## Offset Sideband Locking

- Use EOM as frequency actuator
- EOM already present for PDH
- No modification to cavity
- Increased noise?

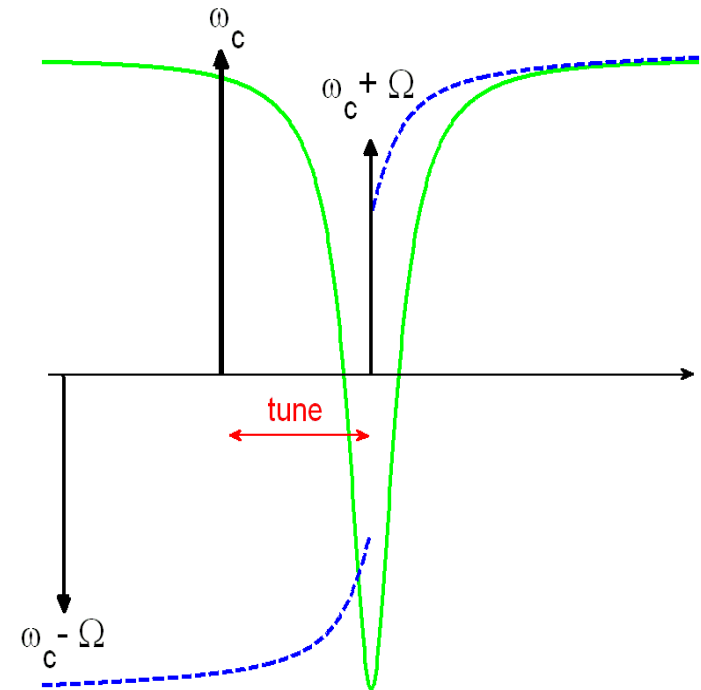


**Concept:** Lock phase-modulation sidebands to cavity resonance and tune central frequency by adjusting modulation frequency.

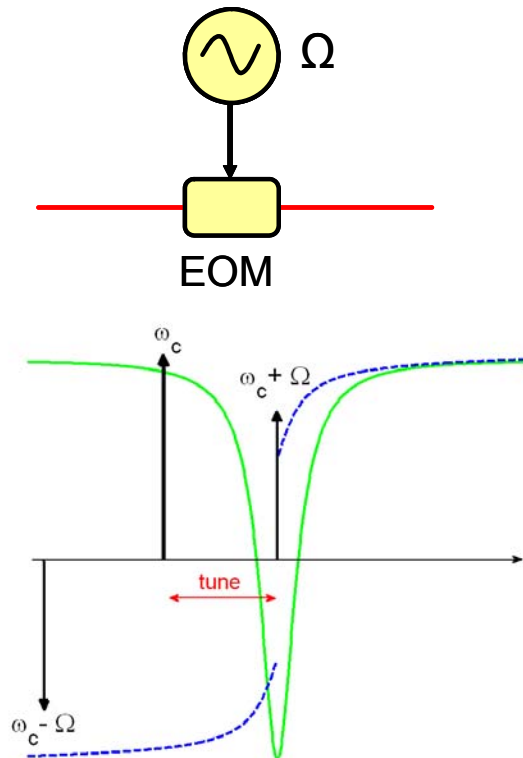
## Normal Pound-Drever-Hall Lock



## Sideband Lock

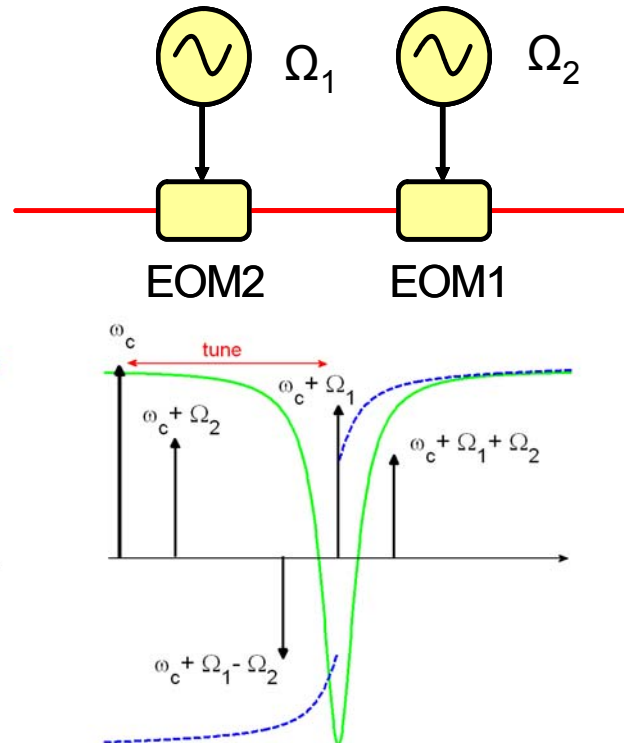


## Single Sideband (SSB)



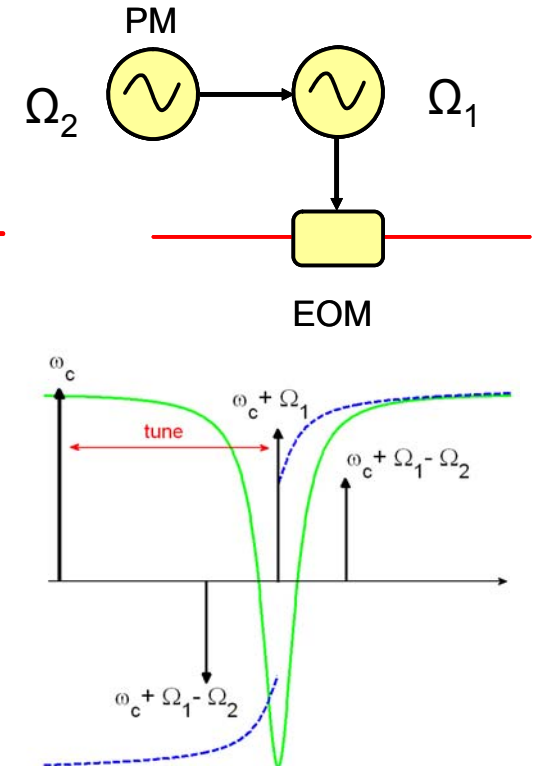
- Simplest to implement
- Some noise coupling due to asymmetry

## Dual Sideband (DSB)

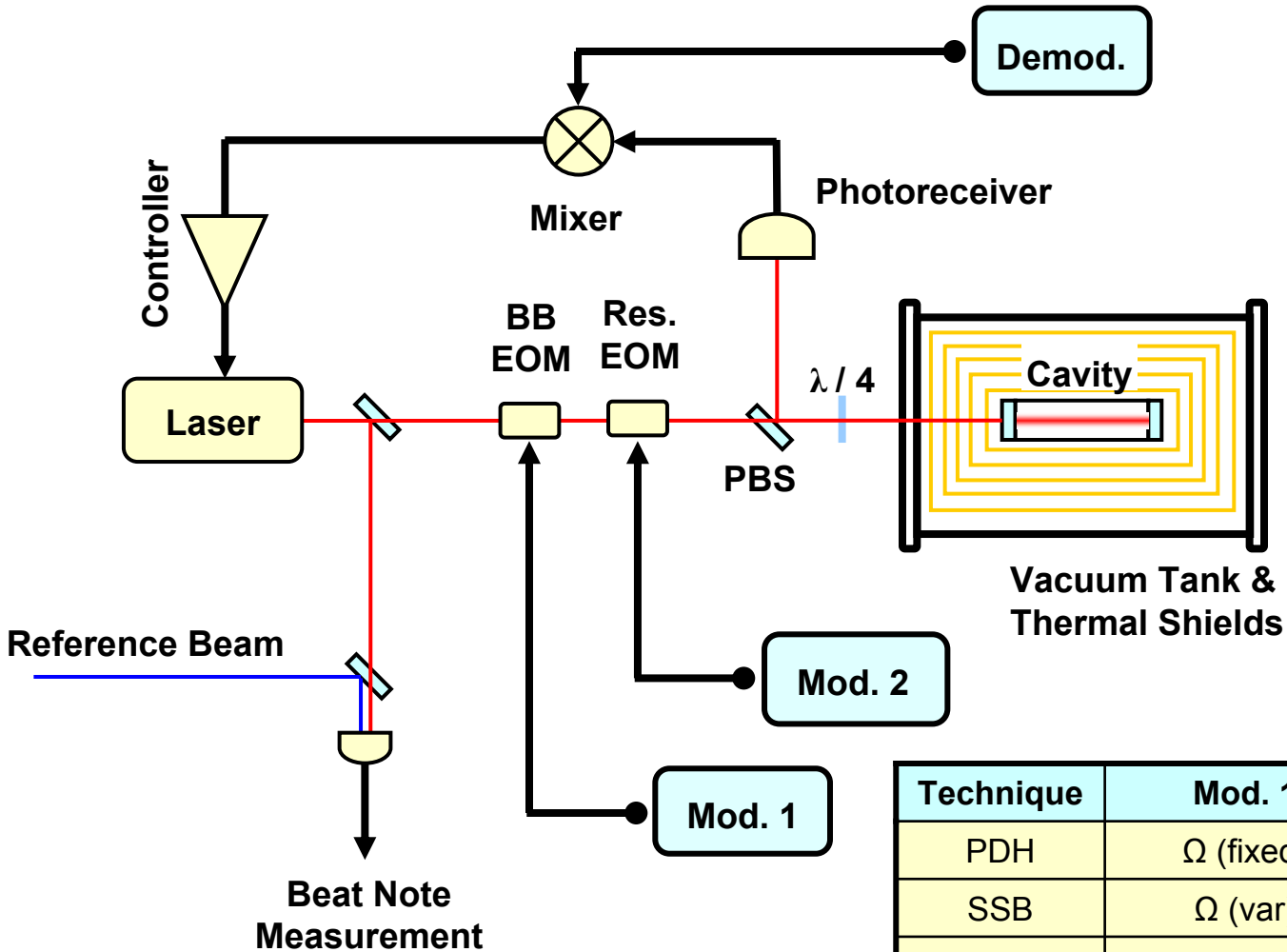


- Restores PDH symmetry
- Complex modulation pattern

## Electronic Sideband (ESB)

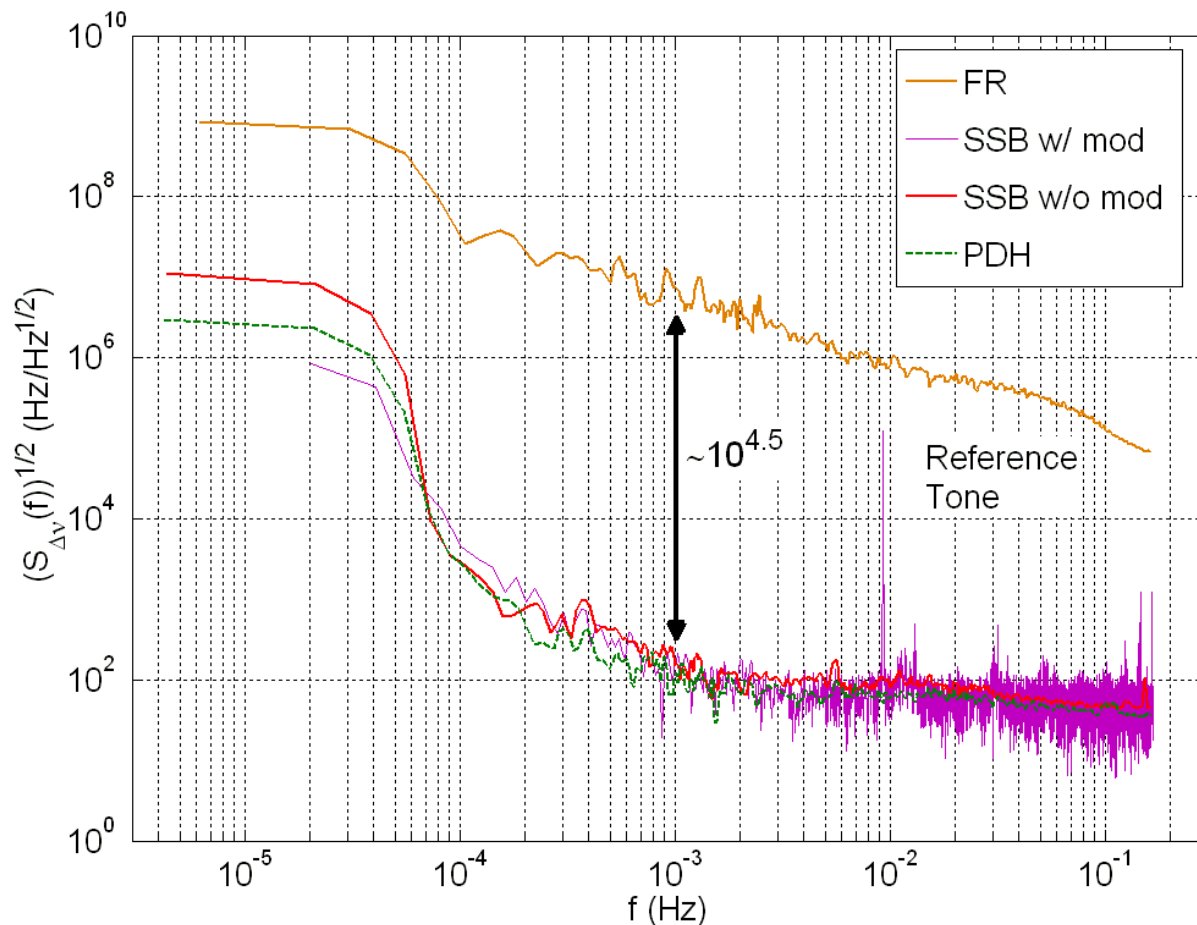


- Simple, symmetric modulation pattern
- Requires phase modulation capability on LO



Technique	Mod. 1	Mod. 2	Demod
PDH	$\Omega$ (fixed)	N/A	$\Omega$ (fixed)
SSB	$\Omega$ (var)	N/A	$\Omega$ (var)
DSB	$\Omega_1$ (var)	$\Omega_2$ (fixed)	$\Omega_2$ (fixed)
ESB	$\Omega_1$ (var) w/ $\Omega_2 \phi M$ (fixed)	N/A	$\Omega_2$ (fixed)

- Standard PDH and SSB locking have identical noise performance.
- Adding modulation tone does not disturb the noise floor.
- Similar results with other sideband locking schemes (DSB & ESB)





# Preliminary Noise Model



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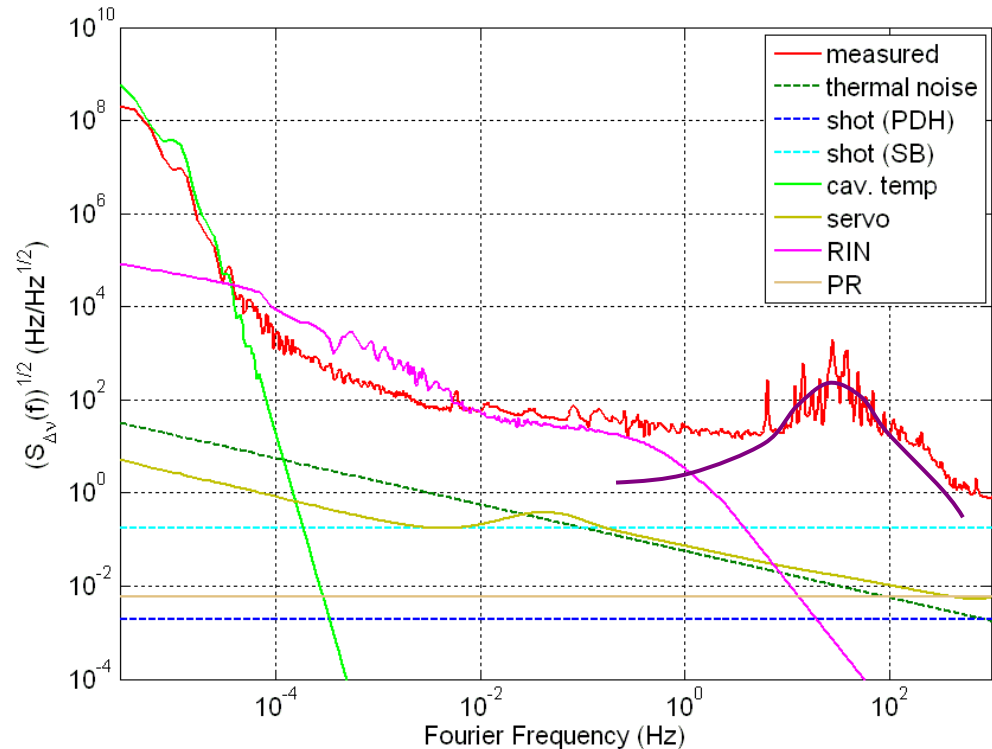
## Fundamental Noise

- Shot noise
- Cavity thermal noise



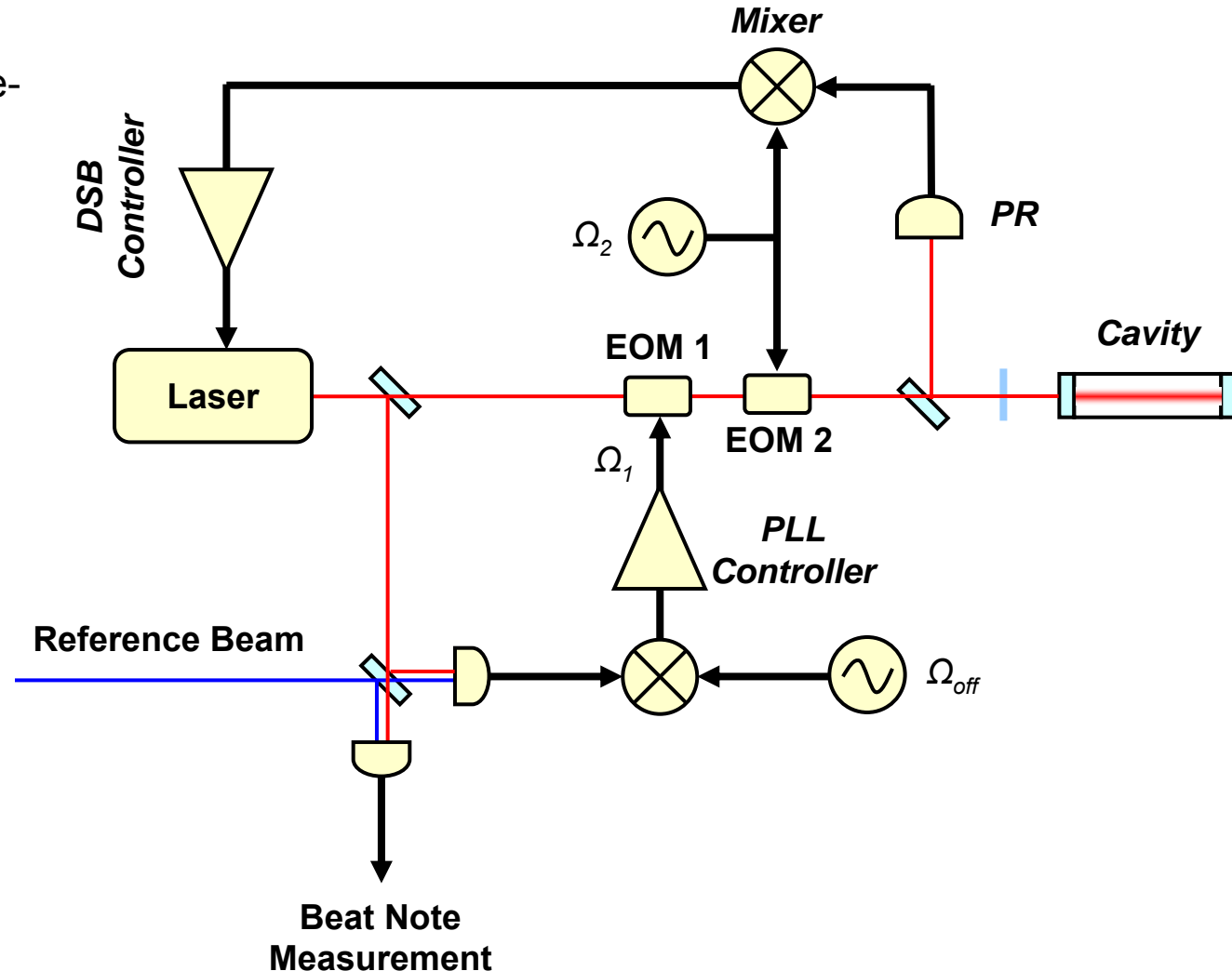
## Technical Noise

- Temperature Fluctuations
- Servo Noise
- Photoreceiver noise
- RIN
  - via RFAM
  - via absorption
- Vibration Noise/Acoustic
- Pointing
- ???

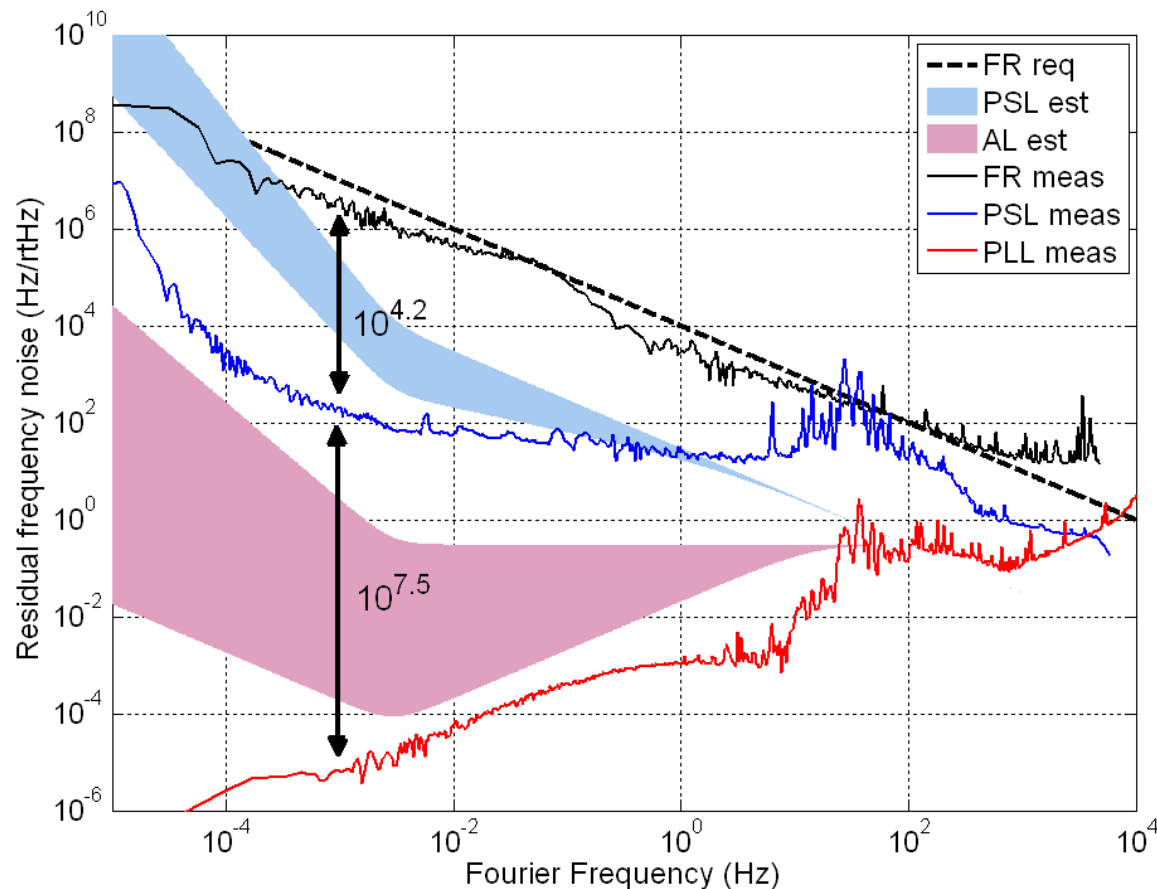




- Use DSB locking as a pre-stabilization stage in a phase-lock loop
- PLL gain required to achieve given noise floor will be reduced by DSB gain.
- Demonstrates dynamic range, bandwidth, and noise of sideband locking system as an actuator.

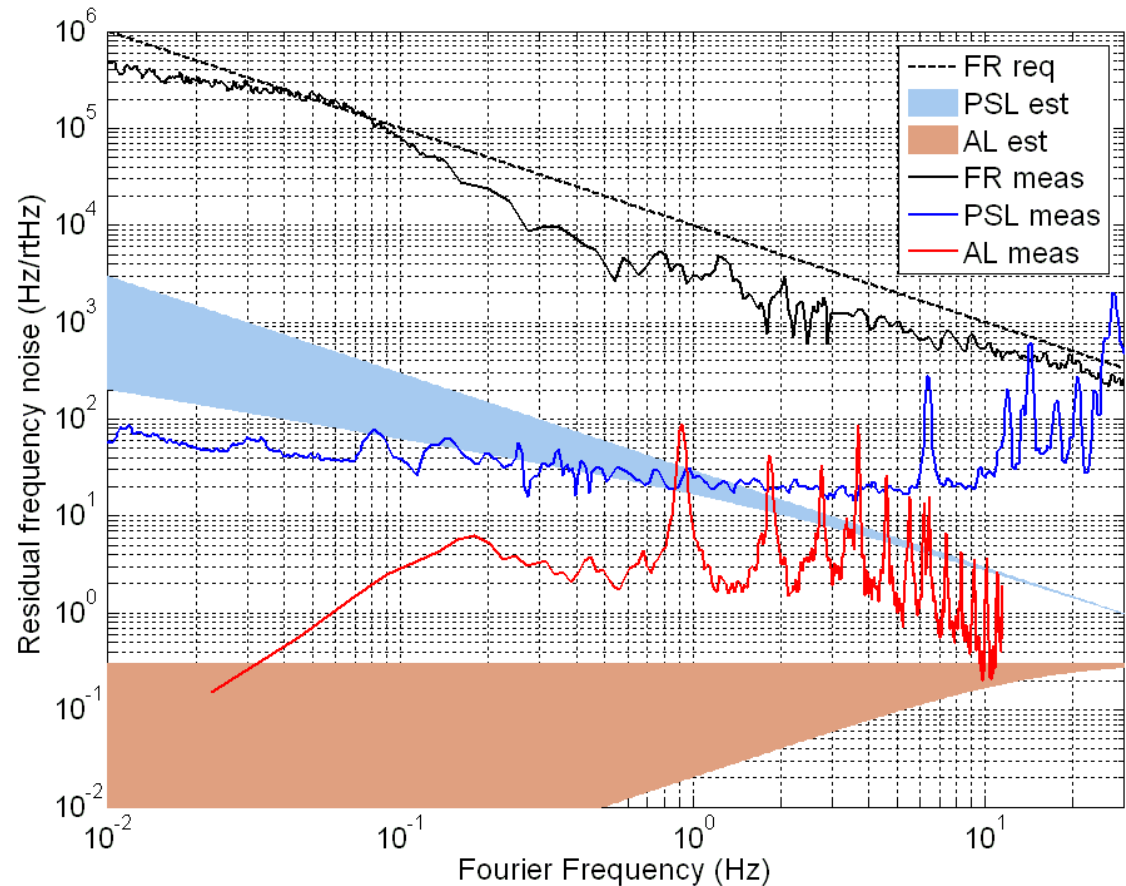


- PSL easily meets requirements below 1Hz
- PSL noise high above 1Hz, believed to be vibration/acoustics & RIN
- PLL noise floor beats arm-locking CBE, indicating actuator noise is sufficiently small for arm-locking



- Simulate 1-s long arm using EPD technique
- Perform single-arm locking in place of PLL
- Observe characteristic arm-locking noise spectrum

See V. Wand's talk  
&  
Y. Yu's poster



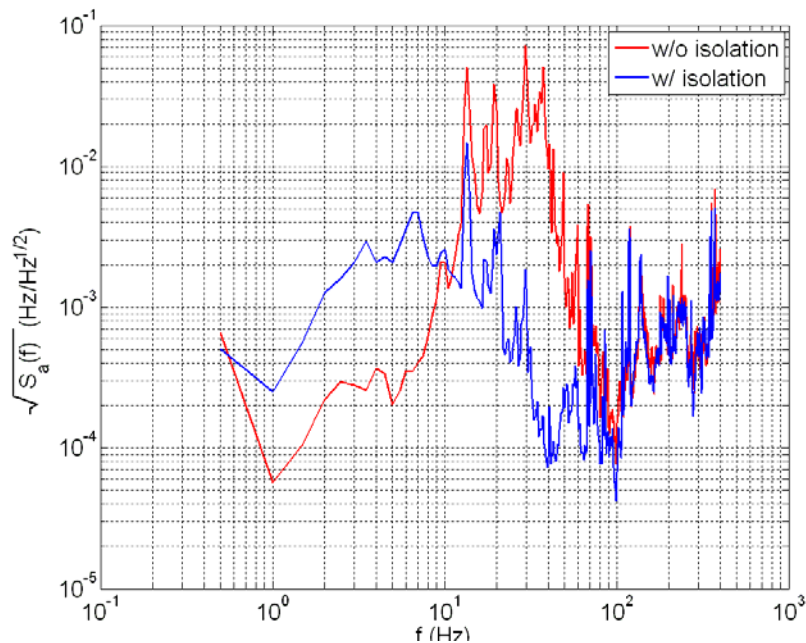
- 🌀 Current LISA Frequency Noise Plan requires frequency-tunable pre-stabilization system
- 🌀 Offset Sideband locking can provide frequency tuning with minimal modification to standard cavity-locking technique
- 🌀 Laboratory results demonstrate that sideband locking can meet the LISA pre-stabilization requirements
- 🌀 Initial demonstration of first two steps to LISA frequency noise suppression.



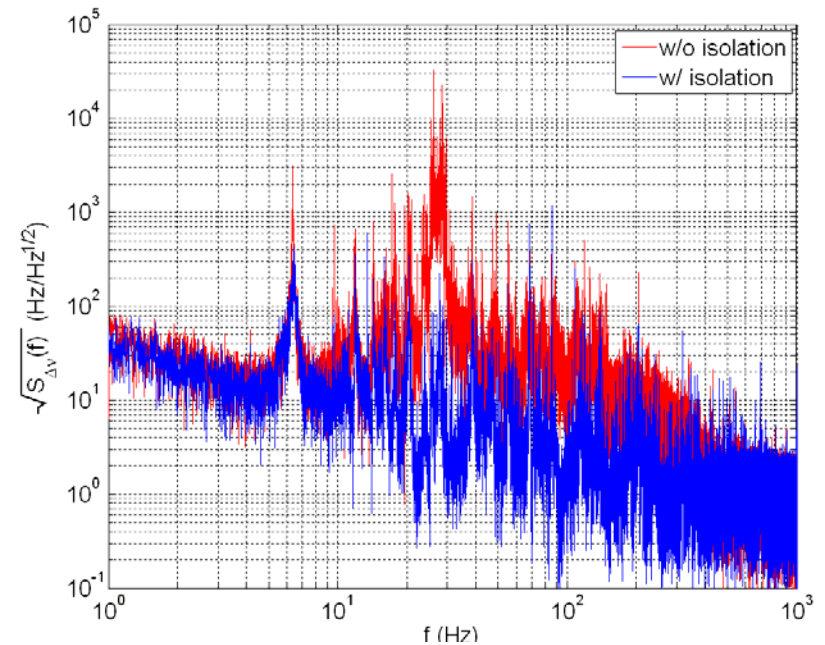
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## *Backup Slides*

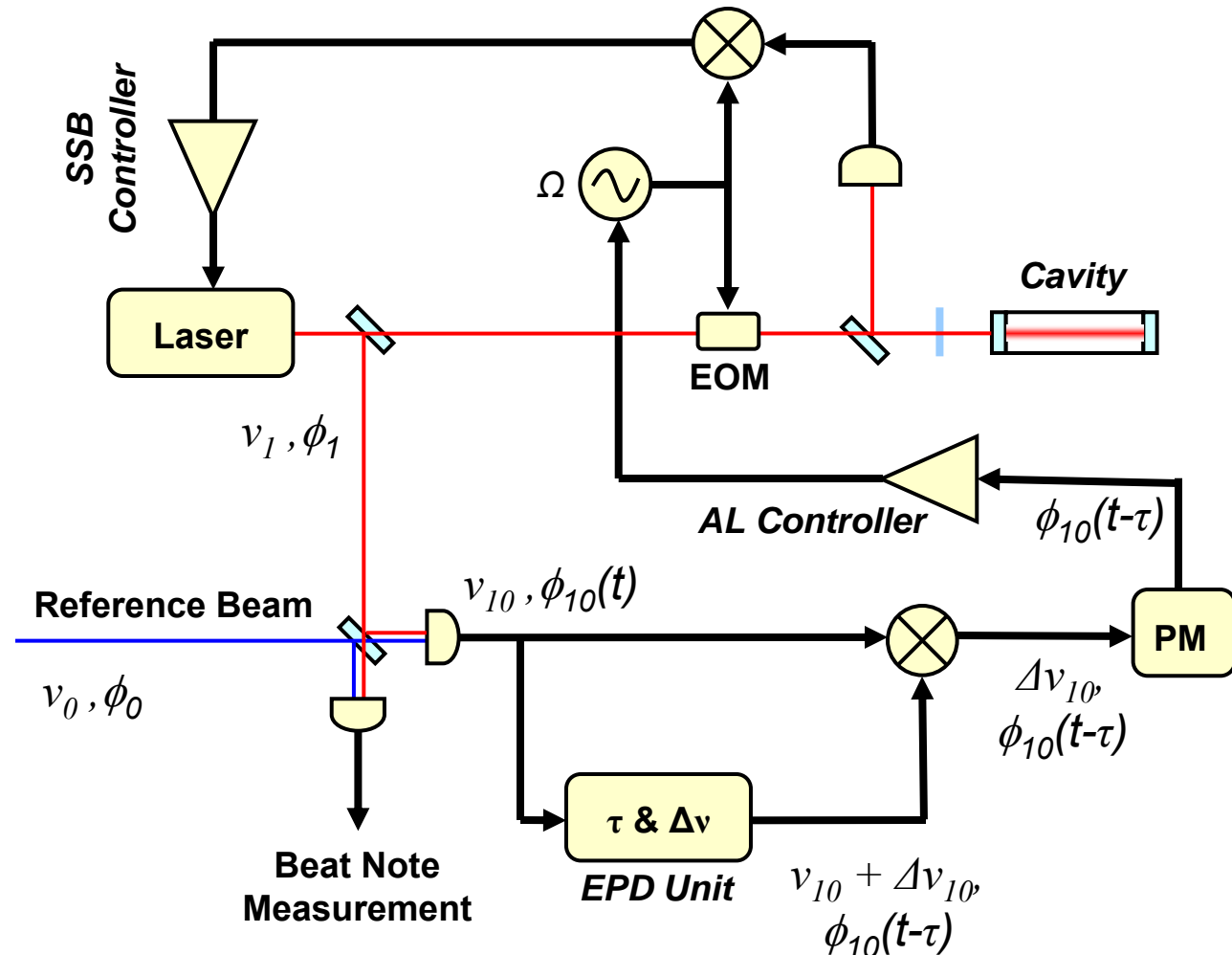
Table Acceleration Noise  
Measured via Accelerometer



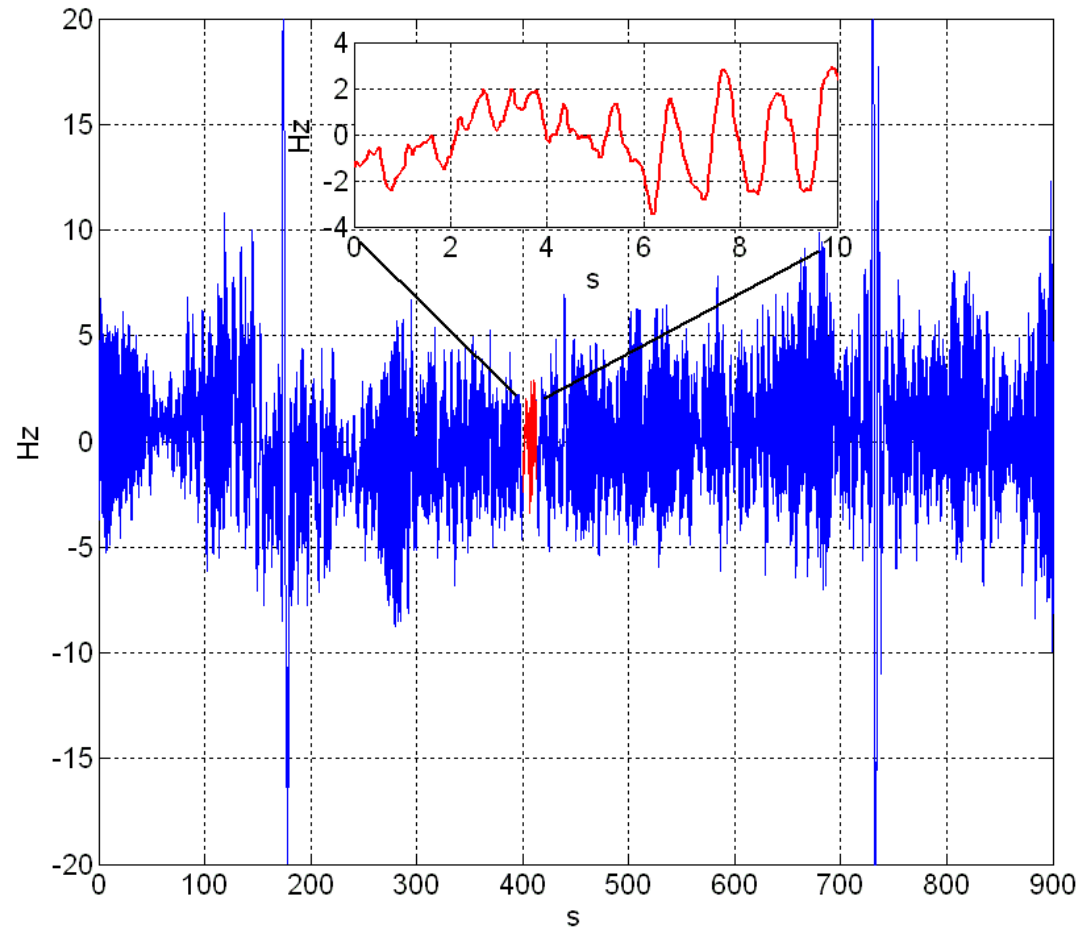
Frequency Noise measured  
via phasemeter



- Extension of PLL example utilizing Electronic Phase Delay (EPD) technique to mimic arm-locking error signal.
- Simulates single-arm locking
- ~1s round-trip delay
- Fixed Doppler shift
- Utilized SSB locking for first test.



- 1 s delay gives noise at 1 Hz multiples
- Pre-stabilization noise is non-stationary, glitches excite arm-locking transients







# LISA

## Arm-Locking Transfer Function

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- Measured noise suppression matches expectations
- ~40dB at 100mHz

